

Topics/Concepts

Tidal torque & dissipation

Synchronous vs. non-synchronous bodies

Kepler's laws, orbital angular momentum and energy, semi-major axis, mean motion

Momentum is conserved but energy isn't

Eccentricity damping, tidal heating

Equations

$$n^2 = \frac{GM}{a^3} \quad E = \frac{-GMm}{2a} \quad L = mna^2\sqrt{1-e^2}$$

$$\text{Tidal torque (non-synchronous)} T = \frac{3}{2} k_2 m^2 G \frac{R_p^5}{a^6} \frac{1}{Q}$$

$$\text{Tidal dissipation (non-synchronous)} \dot{E} = nT$$

$$\text{Tidal dissipation (synchronous)} \dot{E} = \frac{21}{2} \frac{R^5 n^5}{G} e^2 \frac{k_2}{Q} = 7e^2 nT$$

Numbers

$$G = 6.67 \times 10^{-11} \text{ m}^3 \text{ s}^{-2} \text{ kg}^{-1} \quad Q_{\text{solid}} \sim 10^2 \quad Q_{\text{gas}} \sim 10^5$$

Homogeneous, fluid body: $h_2 = 5/2$, $k_2 = 3/2$

References

Murray, C.D., S.F. Dermott, *Solar System Dynamics*, CUP 1999, Chapter 4.

Goldreich, P., S. Soter, Q in the solar system, *Icarus* 5, 375-389, 1966.